



Molecular Physics An International Journal at the Interface Between Chemistry and Physics

ISSN: (Print) (Online) Journal homepage: https://www.tandfonline.com/loi/tmph20

A combined experimental and computational study on the reaction dynamics of the 1-propynyl (CH₃CC, X^2A_1) – propylene (CH₃CHCH₂, X^1A') system: formation of 1,3-dimethylvinylacetylene (CH₃CCCHCHCH₃, X^1A') under single collision conditions

Iakov A. Medvedkov, Anatoliy A. Nikolayev, Chao He, Zhenghai Yang, Alexander M. Mebel & Ralf I. Kaiser

To cite this article: Iakov A. Medvedkov, Anatoliy A. Nikolayev, Chao He, Zhenghai Yang, Alexander M. Mebel & Ralf I. Kaiser (2023): A combined experimental and computational study on the reaction dynamics of the 1-propynyl (CH_3CC , X^2A_1) – propylene (CH_3CHCH_2 , X^1A') system: formation of 1,3-dimethylvinylacetylene ($CH_3CCCHCHCH_3$, X^1A') under single collision conditions, Molecular Physics, DOI: <u>10.1080/00268976.2023.2234509</u>

To link to this article: <u>https://doi.org/10.1080/00268976.2023.2234509</u>

+	View supplementary material 🛽 🖓	Published online: 20 Jul 2023.
	Submit your article to this journal 🕑	View related articles 🕑
CrossMark	View Crossmark data 🗗	

TIM LEE MEMORIAL ISSUE

Taylor & Francis

Check for updates

A combined experimental and computational study on the reaction dynamics of the 1-propynyl (CH₃CC, X^2A_1) – propylene (CH₃CHCH₂, X^1A') system: formation of 1,3-dimethylvinylacetylene (CH₃CCCHCHCH₃, X^1A') under single collision conditions

Iakov A. Medvedkov ^[ba], Anatoliy A. Nikolayev ^{[b,c}, Chao He ^[ba], Zhenghai Yang^a, Alexander M. Mebel^d and Ralf I. Kaiser ^[ba]

^aDepartment of Chemistry, University of Hawaii at Manoa, Honolulu, HI, USA; ^bSamara National Research University, Samara, Russia; ^cLebedev Physical Institute, Samara, Russia; ^dDepartment of Chemistry and Biochemistry, Florida International University, Miami, FL, USA

ABSTRACT

The reaction of the 1-propynyl radical (CH₃CC; X²A₁) with propylene (CH₃CHCH₂; X¹A') was studied in a crossed molecular beam machine at a collision energy of $37 \pm 1 \text{ kJ mol}^{-1}$. Experimental data combined with high-level electronic structure (CCSD(T)-F12/cc-pVTZ-F12//wB97X-D/6-311G(d,p)) and RRKM calculations reveal the reaction mechanism. The overall barrierless and exoergic reaction involves indirect reaction dynamics and commences preferentially with addition of 1-propynyl with its radical centre to the carbon-carbon double bond at the terminal carbon atom of propylene. This work focuses on molecular mass growth process (hydrogen loss channels) although theory suggests methyl loss as a prevalent channel. In these processes, the C_6H_9 collision complexes either emit atomic hydrogen or undergo isomerisation followed by atomic hydrogen loss to preferentially yield the cis/trans isomers of 1,3-dimethylvinylacetylene (2-hexen-4-yne) as the primary product. Analysis of reaction dynamics of 1-propynyl and ethynyl radicals with propylene along with their fractional abundance in deep space suggests formation of methyl- and dimethyl derivatives of vinylacetylene in cold molecular clouds. Once formed they may engage in fundamental molecular mass growth processes via the barrierless Hydrogen Abstraction Vinylacetylene Addition mechanism that leads to the formation of methyl- and dimethylnaphthalenes thus providing a versatile route to methyl-substituted PAHs in interstellar medium.



ARTICLE HISTORY

Received 25 May 2023 Accepted 4 July 2023

KEYWORDS

Reaction dynamics; astrochemistry; alkylated PAHs; HAVA; 1-propynyl

1. Introduction

Up to 20% of the interstellar carbon budget is considered to be locked up as polycyclic aromatic hydrocarbons (PAHs) [1] – molecules consisting of multiple fused aromatic rings – along with their protonated, ionised, (de)hydrogenated, alkylated, and nitrogen-substituted counterparts [2–7]. Their ubiquitous presence has been inferred from the unidentified infrared emission (UIE) bands observed in the range of $3-14 \,\mu\text{m}$ [8] and through the diffuse interstellar bands [9,10], discrete absorption features superimposed on the interstellar extinction curve ranging from the blue part of the visible (400 nm) to the near-infrared (1.2 μ m). PAHs are also suspected to play a significant role in the astrobiological evolution

CONTACT Alexander M. Mebel 🖾 mebela@fiu.edu 💽 Department of Chemistry and Biochemistry, Florida International University, Miami, FL 33199, USA; Ralf I. Kaiser 🖾 ralfk@hawaii.edu 💽 Department of Chemistry, University of Hawaii at Manoa, Honolulu, HI 96822, USA

B Supplemental data for this article can be accessed here. https://doi.org/10.1080/00268976.2023.2234509

of the interstellar medium and are contemplated as key nucleation sites leading to carbonaceous nanoparticles also referred to as interstellar grains [11-14]. Although significant experimental and computational studies have been dedicated toward unravelling the elementary mechanisms synthesising PAHs in circumstellar envelopes [2,15-21], the origin of PAHs in deep space represents still a fundamental paradox of modern astrophysics with typical lifetimes of PAHs of 10⁸ years, but time scales for their injection from carbonrich circumstellar envelopes of 2×10^9 years [22–25]. This discrepancy could be resolved by inferring the existence of hitherto elusive low-temperature routes to a rapid growth of PAHs in molecular clouds to counterbalance their destruction [26]. One of these routes is the rapid, barrierless Hydrogen Abstraction Vinylacetylene Addition (HAVA) mechanism [2,27-32]. HAVA starts with a barrierless formation of a van-der-Waals complex, followed by isomerisation of this complex through an addition of the radical centre of the aromatic radical reactant to the terminal sp²-carbon atom of the vinylacetylene (CH₂CHCCH) molecule leading to a resonantly stabilised free radical (RSFR) intermediate. This RSFR undergoes hydrogen migration from the aromatic ring to the vinylacetylene followed by cyclisation, hydrogen migration, and atomic hydrogen loss along with aromatisation. Compared to competing mechanisms such as Hydrogen Abstraction $-C_2H_2$ (acetylene) Addition (HACA) [17-20] and PhenylAdditionDehydro Cyclization (PAC) [2,21], HAVA has two notable features: (1) the lack of an entrance barrier and (2) a single bimolecular collision event leading de-facto to ring annulation; these characteristics make HAVA an ideal candidate for a low-temperature synthesis of aromatics in molecular clouds.

The formation of alkylated PAHs is even less underaromatics such stood. Alkylated as methylphenanthrene ($C_{15}H_{12}$), methyltriphenylene ($C_{19}H_{14}$), methyl-, dimethyl-, and trimethylnaphtalene ($C_{11}H_{10}$, C₁₂H₁₂, C₁₃H₁₄) were detected in carbonaceous chondrites such as Murchison [33-37], Allende [6,37], Aguas Zarcas, Mukundpura, and Kolang [38,39]. Studies on their formation are contradictory. Speculative works link their origin to aqueous and thermal conditions experienced by the meteorites [40-42], while competing studies suggest that alkylated aromatics were formed prior to their incorporation into meteoritic parent-bodies [37,38,41,43]. Therefore, although meteorites present a record of the chemistry that occurred in the interstellar medium and in the solar nebula, the question 'How alkylated PAHs were formed?' remains open. An answer to this question is important because, at the present time, the 3.4 μ m feature of the UIE [44–46] still cannot be thoroughly accounted for. The C–H stretching in aliphatic group attached to PAHs are believed [47–51] to be responsible for that feature, but the lack of a reasonable mechanism of sustainable formation of alkylated PAHs in interstellar medium places fundamental constraints to this hypothesis [49].

The 1-propynyl radical (CH₃CC; X^2A_1) has received considerable attention for its participation in molecular growth processes to hydrogen deficient hydrocarbons in cold molecular clouds such as toward TMC-1 [52-59]. Albeit 1-propynyl has not been detected in any extraterrestrial environment yet, its potential precursor methylacetylene (propyne; CH₃CCH) holds high fractional abundances of up to 1×10^{-8} in TMC-1 [60,61]. While thermal decomposition [62] of methylacetylene mostly produces the propargyl radical (CH₂CCH), UV photolysis [63-65] can lead to 1-propynyl radical (CH₃CC). This finding reveals that the 1-propynyl radical is likely present in cold molecular clouds. Being higher in energy by 168 kJ mol⁻¹ compared to the propargyl radical [66], it is very reactive and can barrierlessly add to carbon-carbon double and triple bonds of hydrocarbons even at low temperatures [54-58]. This opens up barrier-free pathways to the formation of complex hydrocarbon molecules in cold, carbon-rich environments such as in the atmosphere of Titan [53] and toward molecular clouds like TMC-1 [52,59]. In this context, the 1-propynyl radical can be considered as an alkyl-substituted ethynyl radical (C₂H, $X^2\Sigma^+$). In this case, all 1-propynyl radical (CH₃CC) reactions studied under single collision conditions [54-58] (Figure 1) can be divided into two groups: (1) the methyl group of 1-propynyl acts as a spectator and (2) the methyl group is actively engaged in the reaction dynamics and hence reaction mechanism. The first group includes reactions with acetylene (HCCH; X^2A_1) [54], ethylene (H₂CCH₂; X¹A_{1g}) [57], methylacetylene (CH₃CCH; X¹A₁) [56], 1,3-butadiene (CH₂CH CHCH₂; $X^{1}A_{g}$ [55], and benzene (C₆H₆; $X^{1}A_{1g}$) [58]. These bimolecular reactions lead predominantly to the formation of methyldiacetylene (CH₃CCCCH) [54], 1pentene-3-yne (CH₂CHCCCH₃) [57], dimethyldiacetylene (CH₃CCCCCH₃) [56], toluene (C₆H₅CH₃) [55], and 1-phenyl-propyne (C₆H₅CCCH₃) [58], respectively. The nascent molecules are methyl-substituted analogues of the product of the corresponding ethynyl reaction (Figure 1) [67-74]. On the other hand, the reaction of the 1-propynyl radical with allene (CH₂CCH₂; $X^{1}A_{1}$) [56] can be attributed to the second group. The potential energy surface for bimolecular reaction of 1-propynyl radical with allene includes channels that can lead to eleven distinct reaction products. Pathways to acyclic products are statistically more preferable although two



Figure 1. Elementary reactions of the 1-propynyl radical studied under single-collision conditions. [1] – 1-propynylallene is also formed at a ratio of dimethyldiacetylene versus 1 propynylallene of 9: 0.1. [2] – 1,3-heptadien-5-yne, 5-methylene-1,3-cyclohexadiene, 3-methylene-1-hexen-4-yne are also formed as minor products.

of the three initially formed C_6H_7 collision complexes have two low-energy vibrational modes that increase their life-time and promote rearrangement to the same intermediate. This dynamically narrows all possible pathways to one outcome where fulvene statistically dominates, which was observed experimentally [56]. The analogues reaction of ethynyl with allene only produces acyclic products, i.e. ethynylallene (CHCCHCCH₂) and 1,4-pentadiyne (CHCCH₂CCH) [75].

In this article, we present an experimental and theoretical investigation of the formation of 1,3dimethylvinylacetylene (2-hexen-4-yne) under single collision conditions through the elementary gas phase reaction of the 1-propynyl radical (CH₃CC; X^2A_1) with propylene (CH₃CHCH₂; X^1A') exploiting crossed molecular beams. We also discuss potential reaction pathways of 1, 3-dimethylvinylacetylene (2-hexen-4-yne) to mono- and disubstituted isomers of methylvinylacetylene that adopt the HAVA mechanism leading to distinctly methyl-substituted naphthalene derivatives in the interstellar medium thus expanding our view of how alkylated PAHs may form in deep space.

2. Experimental and computational methods

2.1. Crossed molecular beams

The reaction of the 1-propynyl radical (CH₃CC; X^2A_1) with propylene (CH₃CHCH₂; X^1A') was conducted

under single-collision conditions exploiting a crossed molecular beam machine [76,77]. The experimental setup is discussed in detail in references [54,55,58]; here we will only provide a brief description. A pulsed 1propynyl radical beam was generated via photodissociation (193 nm; 20 mJ pulse⁻¹; 30 Hz) of 1-iodopropyne (CH₃CCI; TCI, 99%+) seeded at a level of 0.5% in helium (99.9999%, Matheson) in the primary source chamber [54,55,58]. This beam was skimmed and velocity selected using a four-slot chopper wheel achieving a peak velocity $v_p = 1721 \pm 38 \text{ m s}^{-1}$ and speed ratio $S = 11.5 \pm 1.8$. This supersonic radical beam crossed a pulsed molecular beam of propylene (Sigma-Aldrich, \geq 99%; $v_p = 840$ \pm 10 m s⁻¹, S = 11.0 \pm 0.2) [78] perpendicularly in the scattering chamber at a mean collision energy of $E_{\rm C} = 37$ \pm 1 kJ mol ⁻¹. The secondary pulsed valve was triggered 86 μ s prior to the primary pulsed valve. The reactively scattered products were detected by a triply differentially pumped universal quadrupole mass spectrometric (QMS) detector under ultra-high vacuum (UHV; 10^{-12} Torr) conditions. The latter is operated in the time-offlight (TOF) mode, i.e. recording the flight time of ions at a distinct mass-to-charge ratio (m/z), and is rotatable within the plane defined by both reactant beams. Neutral products were ionised by electron ionisation at 80 eV (2 mA), filtered according to the selected m/z, and then detected using Daly-type detector [79]. Angular resolved TOF spectra were recorded at discrete laboratory angles,

which were integrated and normalised with respect to the TOF at the centre-of-mass (CM) angle Θ_{CM} to extract the laboratory angular distribution. Up to 7×10^6 TOF spectra were taken in 2.5° steps between $16.25^\circ \leq \Theta \leq 41.25^\circ$ with the 1-propynyl beam defined at $\Theta = 0^\circ$. To gain information on the reaction dynamics, the laboratory angular distribution and TOF spectra were fit utilising a forward convolution routine; this created user-defined CM translational energy (P(E_T)) and angular (T(θ)) flux distributions, which were refined iteratively until a satisfactory fit of the laboratory data was achieved [80,81]. These functions were exploited to develop a flux contour map I(u, θ) \approx P(u) \times T(θ) with the centre-of-mass velocity u [82], which reveals an overall image of the outcome of the reaction.

2.2. Electronic structure calculations

The hybrid density functional theory (DFT) ω B97X-D [83] method in conjunction with the 6-311G(d,p) basis set were utilised for geometry optimisation of various reactants, products, intermediates, and transition states on the C₆H₉ potential energy surface (PES) accessed by the reaction of propynyl radical with propylene and the same ω B97X-D/6-311G(d,p) level of theory was used to compute vibrational frequencies for each stationary structure, with the goal to evaluate zeropoint vibrational energy corrections (ZPE) and to obtain molecular parameters required for rate constant calculations. Next, single-point energies of all optimised structures were refined employing the explicitly correlated



Figure 2. (a) Time-of-flight spectra and (b) laboratory angular distribution recorded at m/z = 79 (C₆H₇⁺) for products formed in the reaction of the 1-propynyl radical (CH₃CC;X²A₁) with propylene (C₃H₆; X¹A'). The circles represent the experimental data and the solid lines the best fit.

couple clusters method with single and double excitations and triple excitations treated perturbatively, CCSD(T)-F12 [84,85], with Dunning's correlation-consistent ccpVTZ-f12 basis set [86]. The overall CCSD(T)-F12/ccpVTZ-f12// ω B97X-D/6-311G(d,p) + ZPE(ω B97X-D/6-311G(d,p)) approach thus used to obtained relative energies of various species on the C₆H₉ PES is expected to be chemically accurate, within 4 kJ mol⁻¹ or better [87]. The GAUSSIAN 09 [88] and MOLPRO 2010 [89] quantum chemistry packages were employed for the DFT and coupled clusters calculations, respectively.

The Rice-Ramsperger-Kassel-Marcus (RRKM) theory method [90-92] was applied to assess energydependent rate constants for all unimolecular reaction steps occurring on the C₆H₉ PES after the entrance bimolecular reaction steps. These computations utilised the relative energies and molecular parameters generated by the electronic structure calculations. The internal energy for each C₆H₉ species was taken as the sum of the collision and chemical activation energy, with the latter being defined as a negative of the relative energy of the particular structure on the PES with respect to the separated $C_3H_3 + C_3H_6$ reactants. The energy-dependent rate constant calculations were carried out using our inhouse code at the zero-pressure limit [56], mimicking the crossed molecular beams and the outer space conditions. Finally, the RRKM-computed rate constants were used to evaluate the product branching ratios within steady-state approximation [56,93].

3. Results

3.1. Laboratory frame

Reactive scattering signal for the reaction of the 1propynyl radical (CH₃CC; 39 amu) with propylene (CH₃CHCH₂; 42 amu) was monitored at mass-to-charge ratios m/z = 80 (C₆H₈⁺), 79 (C₆H₇⁺), and 78 (C₆H₆⁺). The TOF spectra obtained at all three m/z overlap after scaling suggesting that signal from m/z = 80, 79, and 78 originated from the same reaction channel forming the heavy product (C₆H₈, 80 amu) and atomic hydrogen (H, 1 amu) (reaction 1). Signal at m/z = 79 and 78, therefore, arise from dissociative electron ionisation of the C₆H₈ product in the electron ioniser. The background interference at m/z = 66 and 65 prevented detection of a possible methyl loss (reaction 2).

$$CH_{3}CC(39 \text{ amu}) + CH_{3}CHCH_{2}(42 \text{ amu})$$

$$\rightarrow C_{6}H_{8}(80 \text{ amu}) + H(1 \text{ amu}) \qquad (1)$$

$$CH_{3}CC(39 \text{ amu}) + CH_{3}CHCH_{2}(42 \text{ amu})$$

$$\rightarrow C_{6}H_{8}(66 \text{ amu}) + CH_{3}(15 \text{ amu}) \qquad (2)$$

The best signal-to-noise ratio was detected at m/z = 79; hence m/z = 79 was used to collect TOF spectra at discrete angular intervals 2.5° from 16.25° to 41.25° Θ (Figure 2(a)). The resulting TOFs were then normalised with respect to the CM angle to obtain laboratory angular distribution (Figure 2b). Notable features of the laboratory angular distribution include its width of at least 25° and symmetry around the CM angle (28.5 ± 0.3°); these findings propose that the C₆H₈ products were formed via



Figure 3. Centre-of-mass translational energy (a), angular flux distribution (b), and corresponding flux contour map (c) for the formation of C_6H_8 plus atomic hydrogen via the reaction of the 1-propynyl radical (CH₃CC;X²A₁) with propylene (C₃H₆; X¹A'). Grey-filled areas represent regions of acceptable fits.

indirect scattering dynamics through complex formation involving one or more C₆H₉ intermediates [76,82,94].

3.2. Centre-of-mass frame

The laboratory data could be fit with a single channel (reaction 1) with a reaction cross section proportional to $E_{\rm C}^{-(1/3)}$ for entrance-barrierless reactions (Figure 2(a)) dominated by long-range dipole - dipole interactions [95]. The best-fit CM functions are depicted in Figure 3; the grey-filled areas define the limits of the acceptable fits. The high-energy cut-off of the translational energy flux distribution $P(E_T)$ can be determined for those molecules born without internal excitation through energy conservation via $E_{\rm max} = E_{\rm C} - \Delta_{\rm r} G$, where $E_{\rm C}$ and $\Delta_{r}G$ represent the collision energy and the reaction energy, respectively. The derived $P(E_{\rm T})$ reveals an $E_{\rm max}$ of $154 \pm 17 \text{ kJ mol}^{-1}$. Subtracting the collision energy of $37 \pm 1 \text{ kJ mol}^{-1}$, we derived a reaction energy of $117 \pm$ 18 kJ mol^{-1} for the hydrogen loss product. Further, the distribution maximum of the $P(E_{\rm T})$ at 29–52 kJ mol⁻¹ indicates a tight exit translation state in the exit channel and hence a significant electron reorganisation from the decomposing C_6H_9 complex(es) to the final products. The average translation energy of $66 \pm 7 \text{ kJ mol}^{-1}$ suggests that $43 \pm 9\%$ of the total energy is channelled into product translation. Additional information on the reaction dynamics can be obtained by inspecting CM angular distribution $T(\theta)$ (Figure 3(b)). The $T(\theta)$ displays forward-backward symmetry and shows non-zero intensity from 0 to 180°. These findings propose indirect scattering dynamics through C_6H_9 complex(es) and lifetime of C_6H_9 intermediates longer than the(ir) rotational period(s) [95]. Finally, the maximum of the $T(\theta)$ at 90° highlights geometrical constraints of the decomposing complex ('sideways scattering') revealing that the hydrogen atom is eliminated nearly perpendicularly to the plane of decomposing complex and almost parallel to the total angular momentum vector [95,96]. These findings are also supported by the flux contour map (Figure 3(c)), which shows an overall image of the reaction and the scattering process.

4. Discussion

We are combining now our experimental data with electronic structure and statistical calculations to reveal the underlying reaction mechanism(s). The full potential energy surface (PES) (Figure S1-S3), which includes 18 products isomers (p1-p18), 22 intermediates (i1-i22), and 65 transition states, along with results of RRKM calculations (Table S1 and S2) are compiled in the Supporting Information. Reaction energies were calculated with an accuracy of $4 \text{ kJ} \text{ mol}^{-1}$. According to RRKM results, methyl loss channel should dominate in this reaction, however, its experimental detection was hidden by background interference. Notwithstanding, the main objective of this publication is to explore a molecular mass grow process, i.e. the potential formation of p1**p6** along with atomic hydrogen elimination (Figure 4) that together account up to 17% of the total product yield. The energies of formation of 2-methylpent-1-en-3yne (**p1**, -117 kJ mol⁻¹), *cis*-2-hexen-4-yne (**p2**, -117 kJ mol^{-1}), and *trans*-2-hexen-4-yne (p3, -116 kJ mol⁻¹) are in a good agreement with an experimentally derived



Figure 4. Potential energy surface for the bimolecular reaction of the 1-propynyl (CH₃CC; X^2A_1) radical with propylene (CH₃CHCH₂, X^1A') leading to C₆H₈ plus H products calculated at the CCSD(T)-F12/cc-pVTZF12// ω B97X-D/6-311G(d,p) level of theory. Relative energies of various species are shown in kJ mol⁻¹.

reaction energy of -117 ± 18 kJ mol⁻¹. We also cannot exclude the formation of the high-energy isomer 1-methyl-2-vinylidenecyclopropane (**p6**; -28 kJ mol⁻¹) and two conformers of 1-hexen-4-yne (**p4**, -95 kJ mol⁻¹; **p5**, -92 kJ mol⁻¹), since they might be masked in the low energy section of the CM translational energy distribution (Figure 3(a)).

The calculations reveal that the reaction has three barrierless entrance channels via addition of the 1propynyl radical to the double bond of propylene $({}^{\gamma}CH_3 - {}^{\beta}CH = {}^{\alpha}CH_2)$ at the α or β positions leading either to i4/i5 or to intermediate i1, respectively. Intermediate il can undergo migration of the 1-propynyl moiety to the α position via three-carbon ring structures (i2 and i3) to i4 with a barrier of $73 \text{ kJ} \text{ mol}^{-1}$. Alternatively, intermediate **i1** undergoes a ${}^{\gamma}$ CH₃-shift to the ${}^{\alpha}$ C atom forming i6 passing a much higher transition state located only $5 \text{ kJ} \text{ mol}^{-1}$ below the energy of the separated reactants. Isomers i4/i5 can easily be interconverted through the low energy barrier of only 3 kJ mol⁻¹ or form i7/i8, respectively via hydrogen atom shift. Intermediate i6 also can be formed by hydrogen migration from the central CH_2 group of i4/i5. In the next step, i6 can undergo hydrogen migration from the terminal methyl group yielding i7, which can then isomerise to i8 via a barrier of 13 kJ mol⁻¹.

How can the products be formed? p1 can be accessed via hydrogen elimination from il with an exit barrier 143 kJ mol⁻¹, while **p6** is synthesisable from **i2** and **i3**; the unimolecular decompositions of these intermediates have rather lose exit transition states located only 10 kJ mol⁻¹ above the separated products. Note that the formation of p1 and p6 could only be the result of the addition of 1-propynyl to the less sterically accessible β carbon atom of propylene. Pathways leading to formation of **p2-p4** are more diverse and commence with the addition of 1-propynyl to the α position in propylene molecule. Notable, pathways leading to cis/trans 2-hexen-4-yne (p2 and p3) may involve hydrogen loss from i6 – a resonantly stabilised secondary carbon-centered radical which represents the global minimum of this part of the C_6H_9 PES. The computed geometries of these exit transition states leading to p2 and p3 reveal that the hydrogen atom is emitted at angles from 82 to 84° with respect to the rotation plane of the decomposing complexes (Figure 5). This finding is consistent with the experimentally observed sideways scattering depicted in $T(\theta)$ distribution. Therefore, the proposed pathways leading to **p2** and p3 can be accounted for the experimental results.

Which products will dominate? The addition of the electrophilic 1-propynyl radical to the terminal carbon of propylene is favoured due to the higher cone of acceptance (less steric hindrance by the methyl group) and

an enhanced negative charge at α (-0.42) carbon atom as compared to the β (0) position [97]. Combining this addition mechanism with the location of the transition states to isomerisation, **p2/p3** should be the dominant products via reaction sequences (3) and (4). This analysis is supported by our RRKM calculations at collision energy of 37 kJ mol⁻¹ with predicted fractions of 2hexen-4-yne of 67% followed by 1-hexen-4-yne with 30% (Table S1). **p1 + H** only accounts for up to 3%, with calculated statistical branching ratios of high energy isomer **p6** yielding 0.0%. Note that 2-hexen-4yne was also characterised as a main adduct of the hydrogen atom loss channel of the 1-propynyl plus propylene reaction in a slow flow reactor at 300 K via photoionization efficiency (PIE) curves [53]. The aforementioned findings are in

ts i5-p3ts i6-p3Figure 5. Computed geometries of the exit transition states fromi4, i5, and i6 leading to formation of *cis* (p2) and *trans* (p3) isomersof 2-hexen-4-yne. The angle for each departing hydrogen atomis given with respect to the rotation plane of the decomposingcomplex.





Figure 6. Reaction pathways leading to formation of methyl-substituted naphthalenes.

line with the results of the reaction of the D1-ethynyl radical (C₂D; X² Σ^+) with propylene (CH₃CHCH₂; X¹A[']) under single collision conditions [78]. In the ethynyl system, *cis/trans*-3-penten-1-yne ((HCC)CH = CH(CH₃), 3-methylvinylacetylene) – the mono-methyl substituted homologous counterpart – represents the dominating product. This reveals that the methyl group of the 1propynyl radical acts solely as a spectator in the reaction with propylene.

$$[i1 \rightarrow i2 \rightarrow i3 \rightarrow i4 \rightarrow i6]/[i1 \rightarrow i6] \rightarrow p2 + H \eqref{3}$$

$$\begin{array}{l} \mathrm{CH}_{3}\mathrm{CC} + \mathrm{CH}_{3}\mathrm{CH}\mathrm{CH}_{2} \rightarrow \mathbf{i5}/\\ [\mathbf{i5} \rightarrow \mathbf{i6}]/[\mathbf{i1} \rightarrow \mathbf{i2} \rightarrow \mathbf{i3} \rightarrow \mathbf{i4} \rightarrow \mathbf{i5}]/\\ [\mathbf{i1} \rightarrow \mathbf{i2} \rightarrow \mathbf{i3} \rightarrow \mathbf{i4} \rightarrow \mathbf{i5} \rightarrow \mathbf{i6}]/[\mathbf{i1} \rightarrow \mathbf{i6}]\\ \rightarrow \mathbf{p3} + \mathrm{H} \end{array} \tag{4}$$

$$\begin{array}{l} \mbox{CH}_3\mbox{CC} + \mbox{CH}_3\mbox{CH}_2 \rightarrow i4 / \\ [i4 \rightarrow i6] / [i1 \rightarrow i2 \rightarrow i3 \rightarrow i4] / \end{array}$$

5. Conclusion

We conducted the crossed molecular beam reaction of the 1-propynyl radical (CH₃CC; X^2A_1) with propylene

(CH₃CHCH₂; X^1A') at a collision energy of $37 \pm 1 \text{ kJ}$ mol⁻¹ and merged the experimental data with electronic structure and RRKM calculations. The overall barrierless and exoergic reaction involves indirect reaction dynamics and commences preferentially with the addition of the 1-propynyl radical with its radical centre to the carbon-carbon double bond at the terminal ($^{\alpha}$ CH₂) carbon atom of propylene. RRKM results suggest that methyl loss channel should dominate in this reaction with branching ratio exceeding 80%, but its experimental detection was prevented by background interference. In this work we have focused on H-loss channels to explore a molecular mass grow processes accounted up to 17% of the total product yield. In these processes the C₆H₉ collision complexes either emit atomic hydrogen or undergo isomerisation via 1,2-H shift from ${}^{\alpha}CH_2$ group to ${}^{\beta}C$ -position followed by atomic hydrogen loss to preferentially yield the cis/trans isomers of 1,3-dimethylvinylacetylene (2-hexen-4-yne) as the primary product under single collision conditions. The methyl group of the 1propynyl radical reactant remains a spectator throughout the reaction. Considering that the reaction has no entrance barrier, is exoergic, and all transition states involved are located below the energy of the separated reactants, the formation of 1,3-dimethylvinylacetylene (CH₃CCCHCHCH₃) via the reaction of the 1-propynyl radical with propylene represents a powerful molecular growth process to form a dimethylsubstituted vinylacetylene product - 2-hexen-4-yne - via a single collision event.

In deep space, propylene was detected toward TMC-1 with fractional abundances of 2 $\times 10^{-9}$ [98], while 1-propynyl is highly likely to be present in TMC-1 since its precursor methylacetylene holds a high fraction in TMC-1 up to 1×10^{-8} [60,61]. Therefore, we may predict that 1,3-dimethylvinylacetylene can be formed easily in TMC-1. Analogues barrierless reactions of the ethynyl radical with fractional abundances of 5 $\times 10^{-9}$ in OMC-1 [99] with propylene [78], and of 1-propynyl with ethylene (CRL 618 [100]) [54] may operate in cold molecular clouds. Once methyl- and dimethyl derivatives of vinylacetylene have been formed, these hydrogen deficient reactants may engage in fundamental molecular mass growth processes upon reactions with phenyl radials (C₆H₅) [29] and tolyl radicals (CH₃C₆H₅) [31] via the barrierless HAVA mechanism (Figure 6). These processes can then yield to the formation of methyl and dimethyl naphthalenes thus providing a versatile route to methyl substituted aromatics and leading to a better understanding of the question 'How alkylated PAHs form in interstellar medium?'.

Disclosure statement

No potential conflict of interest was reported by the author(s).

Funding

This work was supported by the U.S. Department of Energy, Basic Energy Sciences (grant nos. DE-FG02-03ER15411 and DE-FG02-04ER15570) to the University of Hawaii and Florida International University, respectively. Ab initio calculations in Samara were supported by the Ministry of Science and Higher Education of the Russian Federation under grant number 075-15-2021-597.

ORCID

Iakov A. Medvedkov D http://orcid.org/0000-0003-0672-2090 Anatoliy A. Nikolayev D http://orcid.org/0000-0002-1733-3704

Ralf I. Kaiser b http://orcid.org/0000-0002-7233-7206 Chao He b http://orcid.org/0000-0001-9351-5684

References

- [1] L. d'Hendecourt and P. Ehrenfreund, Adv. Space Res.
 19 (7), 1023–1032 (1997). doi:10.1016/S0273-1177(97) 00349-9.
- [2] R.I. Kaiser and N. Hansen, J. Phys. Chem. A. 125 (18), 3826–3840 (2021). doi:10.1021/acs.jpca.1c00606.
- [3] E.F. van Dishoeck, Publ. Astron. Soc. Pac. 112 (768), 286–287 (2000). doi:10.1086/316509.
- [4] H. Naraoka, A. Shimoyama and K. Harada, Earth Planet. Sci. Lett. 184 (1), 1–7 (2000). doi:10.1016/S0012-821X (00)00316-2.
- [5] N. Balucani, O. Asvany, Y. Osamura, L.C.L. Huang, Y.T. Lee and R.I. Kaiser, Planet. Space Sci. 48 (5), 447–462 (2000). doi:10.1016/S0032-0633(00)00018-0.
- [6] L. Becker and T.E. Bunch, Meteorit. Planet. Sci.
 32 (4), 479–487 (1997). doi:10.1111/j.1945-5100.1997. tb01292.x.
- M.C. McCarthy and B.A. McGuire, J. Phys. Chem. A. 125 (16), 3231–3243 (2021). doi:10.1021/acs.jpca.1c00129.
- [8] A.M. Ricks, G.E. Douberly and M.A. Duncan, Astrophys. J. **702** (1), 301–306 (2009). doi:10.1088/0004-637X/702/1/301.
- [9] F. Salama, G.A. Galazutdinov, J. Krełowski, L.J. Allamandola and F.A. Musaev, Astrophys. J. **526** (1), 265–273 (1999). doi:10.1086/307978.
- [10] W.W. Duley, Faraday Discuss. 133, 415 (2006). doi:10. 1039/b516323d.
- [11] J.L. Puget and A. Léger, Annu. Rev. Astron. Astrophys. 27
 (1), 161–198 (1989). doi:10.1146/annurev.aa.27.090189.
 001113.
- [12] P. Ehrenfreund and M.A. Sephton, Faraday Discuss. 133, 277 (2006). doi:10.1039/b517676j.
- [13] A.G.G.M. Tielens, Annu. Rev. Astron. Astrophys. 46 (1), 289–337 (2008). doi:10.1146/annurev.astro.46.060407. 145211.
- [14] L.M. Ziurys, Proc. Natl. Acad. Sci. 103 (33), 12274–9 (2006). doi:10.1073/pnas.0602277103.

- [15] E.R. Micelotta, A.P. Jones and A.G.G.M. Tielens, Astron. Astrophys. 510, A36 (2010). doi:10.1051/0004-6361/20 0911682.
- [16] E.R. Micelotta, A.P. Jones and A.G.G.M. Tielens, Astron. Astrophys. 510, A37 (2010). doi:10.1051/0004-6361/20 0911683.
- [17] T. Yang, R.I. Kaiser, T.P. Troy, B. Xu, O. Kostko, M. Ahmed, A.M. Mebel, M.V. Zagidullin and V.N. Azyazov, Angew. Chem. Int. Ed. 56 (16), 4515–4519 (2017). doi:10.1002/anie.201701259.
- [18] D.S.N. Parker, R.I. Kaiser, T.P. Troy and M. Ahmed, Angew. Chem. Int. Ed. 53 (30), 7740–7744 (2014). doi:10.1002/anie.201404537.
- [19] L. Zhao, R.I. Kaiser, B. Xu, U. Ablikim, M. Ahmed, D. Joshi, G. Veber, F.R. Fischer and A.M. Mebel, Nat. Astron. 2 (5), 413–419 (2018). doi:10.1038/s41550-018-0399-y.
- [20] T. Yang, T.P. Troy, B. Xu, O. Kostko, M. Ahmed, A.M. Mebel and R.I. Kaiser, Angew. Chem. Int. Ed. 55 (48), 14983–7 (2016). doi:10.1002/anie.201607509.
- [21] L. Zhao, M.B. Prendergast, R.I. Kaiser, B. Xu, U. Ablikim, M. Ahmed, B. Sun, Y. Chen, A.H.H. Chang, R.K. Mohamed and F.R. Fischer, Angew. Chem. Int. Ed. 58 (48), 17442–17450 (2019). doi:10.1002/anie.201909876.
- [22] M. Frenklach and E.D. Feigelson, Astrophys. J. 341, 372 (1989). doi:10.1086/167501.
- [23] M. Cohen, A.G.G.M. Tielens and J.D. Bregman, Astrophys. J. 344, L13 (1989). doi:10.1086/185519.
- [24] I. Cherchneff, J.R. Barker and A.G.G.M. Tielens, Astrophys. J. 401, 269 (1992). doi:10.1086/172059.
- [25] I. Cherchneff, EAS Publ. Ser. 46, 177–189 (2011). doi:10.1051/eas/1146019.
- [26] R.M. Mahfouz, M. Sauer, S.T. Atwa, R.I. Kaiser and K. Roessler, Nucl. Instrum. Methods Phys. Res. Sect. B Beam Interact. Mater. At. 65 (1–4), 447–451 (1992). doi:10.1016/0168-583X(92)95083-4.
- [27] L. Zhao, R.I. Kaiser, B. Xu, U. Ablikim, M. Ahmed, M.M. Evseev, E.K. Bashkirov, V.N. Azyazov and A.M. Mebel, Nat. Astron. 2 (12), 973–979 (2018). doi:10.1038/s415 50-018-0585-y.
- [28] B. Shukla and M. Koshi, Combust. Flame. 159 (12), 3589–3596 (2012). doi:10.1016/j.combustflame.2012. 08.007.
- [29] D.S.N. Parker, F. Zhang, Y.S. Kim, R.I. Kaiser, A. Landera, V.V. Kislov, A.M. Mebel and A.G.G.M. Tielens, Proc. Natl. Acad. Sci. **109** (1), 53–58 (2012). doi:10.1073/pnas.1113827108.
- [30] T. Yang, L. Muzangwa, R.I. Kaiser, A. Jamal and K. Morokuma, Phys. Chem. Chem. Phys. 17 (33), 21564–21575 (2015). doi:10.1039/C5CP03285G.
- [31] D.S.N. Parker, B.B. Dangi, R.I. Kaiser, A. Jamal, M.N. Ryazantsev, K. Morokuma, A. Korte and W. Sander, J. Phys. Chem. A. **118** (15), 2709–2718 (2014). doi:10.1021 /jp501210d.
- [32] R.I. Kaiser, D.S.N. Parker and A.M. Mebel, Annu. Rev. Phys. Chem. 66 (1), 43–67 (2015). doi:10.1146/annurevphyschem-040214-121502.
- [33] R. Hayatsu, S. Matsuoka, R.G. Scott, M.H. Studier and E. Anders, Geochim. Cosmochim. Acta. 41 (9), 1325–1339 (1977). doi:10.1016/0016-7037(77)90076-X.
- [34] J.R. Cronin and S. Chang, in *The Chemistry of Life's Ori*gins, edited by J.M. Greenberg, C.X. Mendoza-Gómez, V.

Pirronello (Springer Netherlands, Dordrecht, 1993). pp. 209–258.

- [35] J.M. Hayes, Geochim. Cosmochim. Acta. 31 (9), 1395–1440 (1967). doi:10.1016/0016-7037(67)90019-1.
- [36] M.P. Callahan, A. Abo-Riziq, B. Crews, L. Grace and M.S. de Vries, Spectrochim. Acta. A. Mol. Biomol. Spectrosc. 71 (4), 1492–1495 (2008). doi:10.1016/j.saa.2008.05. 005.
- [37] M.A. Sephton, in *Treatise on Geochemistry* (Elsevier, Amsterdam, 2014). pp. 1–31.
- [38] M. Lecasble, L. Remusat, J.-C. Viennet, B. Laurent and S. Bernard, Geochim. Cosmochim. Acta. 335, 243–255 (2022). doi:10.1016/j.gca.2022.08.039.
- [39] M.S. Kalpana, E.V.S.S.K. Babu, D. Mani, R.P. Tripathi and N. Bhandari, Planet. Space Sci. 198, 105177 (2021). doi:10.1016/j.pss.2021.105177.
- [40] J.E. Elsila, N.P. de Leon, P.R. Buseck and R.N. Zare, Geochim. Cosmochim. Acta. 69 (5), 1349–1357 (2005). doi:10.1016/j.gca.2004.09.009.
- [41] M.A. Sephton, C.T. Pillinger and I. Gilmour, Geochim. Cosmochim. Acta. 64 (2), 321–328 (2000). doi:10.1016/ S0016-7037(99)00282-3.
- [42] S. Messenger, S. Amari, X. Gao, R.M. Walker, S.J. Clemett, X.D.F. Chillier, R.N. Zare and R.S. Lewis, Astrophys. J. 502 (1), 284–295 (1998). doi:10.1086/305 874.
- [43] F.L. Plows, J.E. Elsila, R.N. Zare and P.R. Buseck, Geochim. Cosmochim. Acta. 67 (7), 1429–1436 (2003). doi:10.1016/S0016-7037(02)01277-2.
- [44] B.J. Hrivnak, T.R. Geballe and S. Kwok, Astrophys. J. 662
 (2), 1059–1066 (2007). doi:10.1086/518109.
- [45] T.R. Geballe, A.G.G.M. Tielens, S. Kwok and B.J. Hrivnak, Astrophys. J. 387, L89 (1992). doi:10.1086/186312.
- [46] S. Kwok and Y. Zhang, Nature. 479 (7371), 80–83 (2011). doi:10.1038/nature10542.
- [47] M. Jourdain de Muizon, L.B. D'Hendecourt and T.R. Geballe, Astron. Astrophys. 235, 367 (1990).
- [48] M. Jourdain de Muizon, P. Cox and J. Lequeux, Astron. Astrophys. Suppl. Ser. 83, 337–355 (1990).
- [49] S. Kwok, Astrophys. Space Sci. 367 (2), 16 (2022). doi:10.1007/s10509-022-04045-6.
- [50] X.J. Yang, R. Glaser, A. Li and J.X. Zhong, Astrophys.
 J. 776 (2), 110 (2013). doi:10.1088/0004-637X/776/2/ 110.
- [51] A. Li and B.T. Draine, Astrophys. J. 760 (2), L35 (2012). doi:10.1088/2041-8205/760/2/L35.
- [52] S. Doddipatla, G.R. Galimova, H. Wei, A.M. Thomas, C. He, Z. Yang, A.N. Morozov, C.N. Shingledecker, A.M. Mebel and R.I. Kaiser, Sci. Adv. 7 (1), eabd4044 (2021). doi:10.1126/sciadv.abd4044.
- [53] B.B. Kirk, J.D. Savee, A.J. Trevitt, D.L. Osborn and K.R. Wilson, Phys. Chem. Chem. Phys. 17 (32), 20754–20764 (2015). doi:10.1039/C5CP02589C.
- [54] A.M. Thomas, L. Zhao, C. He, A.M. Mebel and R.I. Kaiser, J. Phys. Chem. A. **122** (33), 6663–6672 (2018). doi:10.1021/acs.jpca.8b05530.
- [55] A.M. Thomas, C. He, L. Zhao, G.R. Galimova, A.M. Mebel and R.I. Kaiser, J. Phys. Chem. A. **123** (19), 4104–4118 (2019). doi:10.1021/acs.jpca.9b00092.
- [56] C. He, L. Zhao, A.M. Thomas, A.N. Morozov, A.M. Mebel and R.I. Kaiser, J. Phys. Chem. A. **123** (26), 5446–5462 (2019). doi:10.1021/acs.jpca.9b03746.

- [57] C. He, L. Zhao, A.M. Thomas, G.R. Galimova, A.M. Mebel and R.I. Kaiser, Phys. Chem. Chem. Phys. 21 (40), 22308–22319 (2019). doi:10.1039/C9CP04073K.
- [58] A.M. Thomas, S. Doddipatla, R.I. Kaiser, G.R. Galimova and A.M. Mebel, Sci. Rep. 9 (1), 17595 (2019). doi:10.1038/s41598-019-53987-5.
- [59] A.M. Burkhardt, K. Long Kelvin Lee, P. Bryan Changala, C.N. Shingledecker, I.R. Cooke, R.A. Loomis, H. Wei, S.B. Charnley, E. Herbst, M.C. McCarthy and B.A. McGuire, Astrophys. J. Lett. **913** (2), L18 (2021). doi:10.3847/2041-8213/abfd3a.
- [60] V.V. Guzmán, J. Pety, P. Gratier, J.R. Goicoechea, M. Gerin, E. Roueff, F. Le Petit and J. Le Bourlot, Faraday Discuss. 168, 103–127 (2014). doi:10.1039/C3FD0 0114H.
- [61] P. Gratier, L. Majumdar, M. Ohishi, E. Roueff, J.C. Loison, K.M. Hickson and V. Wakelam, Astrophys. J. Suppl. Ser. 225 (2), 25 (2016). doi:10.3847/0067-0049/225/ 2/25.
- [62] M.N. Ryazantsev, A. Jamal, S. Maeda and K. Morokuma, Phys. Chem. Chem. Phys. 17 (41), 27789–27805 (2015). doi:10.1039/C5CP04329H.
- [63] S. Harich, J.J. Lin, Y.T. Lee and X. Yang, J. Chem. Phys. 112 (15), 6656–6665 (2000). doi:10.1063/1.481316.
- [64] W. Sun, K. Yokoyama, J.C. Robinson, A.G. Suits and D.M. Neumark, J. Chem. Phys. **110** (9), 4363–4368 (1999). doi:10.1063/1.478318.
- [65] Y. Ganot, S. Rosenwaks and I. Bar, J. Chem. Phys. 120 (18), 8600–8607 (2004). doi:10.1063/1.1698679.
- [66] P. Maksyutenko, F. Zhang, X. Gu and R.I. Kaiser, Phys. Chem. Chem. Phys. 13 (1), 240–252 (2011). doi:10.1039/C0CP01529F.
- [67] R.I. Kaiser, ESA. Sp. Publ. 496, 145-153 (2001).
- [68] R.I. Kaiser and N. Balucani, Int. J. Astrobiol. 1 (1), 15–23 (2002). doi:10.1017/S1473550402001015.
- [69] F. Zhang, Y.S. Kim, R.I. Kaiser, S.P. Krishtal and A.M. Mebel, J. Phys. Chem. A. 113 (42), 11167–11173 (2009). doi:10.1021/jp9032595.
- [70] B.M. Jones, F. Zhang, R.I. Kaiser, A. Jamal, A.M. Mebel, M.A. Cordiner and S.B. Charnley, Proc. Natl. Acad. Sci. 108 (2), 452–457 (2011). doi:10.1073/pnas.1012468 108.
- [71] B. Jones, F. Zhang, P. Maksyutenko, A.M. Mebel and R.I. Kaiser, J. Phys. Chem. A. **114** (16), 5256–5262 (2010). doi:10.1021/jp912054p.
- [72] F. Stahl, P. von Ragué Schleyer, H.F. Bettinger, R.I. Kaiser,
 Y.T. Lee and H.F. Schaefer III, J. Chem. Phys. 114 (8),
 3476–3487 (2001). doi:10.1063/1.1331360.
- [73] F. Stahl, P.v.R. Schleyer, H.F. Schaefer III and R.I. Kaiser, Planet. Space Sci. 50 (7), 685–692 (2002). doi:10.1016/S0032-0633(02)00014-4.
- [74] R.I. Kaiser, C.C. Chiong, O. Asvany, Y.T. Lee, F. Stahl, P.R. Schleyer and H.F. Schaefer III, J. Chem. Phys. **114** (8), 3488–3496 (2001). doi:10.1063/1.1330233.
- [75] F. Zhang, S. Kim and R.I. Kaiser, Phys. Chem. Chem. Phys. 11 (23), 4707 (2009). doi:10.1039/b822366a.
- [76] R.I. Kaiser, P. Maksyutenko, C. Ennis, F. Zhang, X. Gu, S.P. Krishtal, A.M. Mebel, O. Kostko and M. Ahmed, Faraday Discuss. 147, 429 (2010). doi:10.1039/c003 599h.
- [77] Y. Guo, X. Gu, E. Kawamura and R.I. Kaiser, Rev. Sci. Instrum. 77 (3), 034701 (2006). doi:10.1063/1.2173074.

- [78] S.J. Goettl, C. He, D. Paul, A.A. Nikolayev, V.N. Azyazov, A.M. Mebel and R.I. Kaiser, J. Phys. Chem. A. **126** (11), 1889–1898 (2022). doi:10.1021/acs.jpca.2c00297.
- [79] N.R. Daly, Rev. Sci. Instrum. 31 (3), 264–267 (1960). doi:10.1063/1.1716953.
- [80] M.F. Vernon, Molecular beam scattering. Ph.D. dissertation, University of California, Berkeley, CA, 1983.
- [81] P.S. Weiss, Reaction dynamics of electronically excited alkali atoms with simple molecules. Ph.D. dissertation, University of California, Berkeley, CA, 1986.
- [82] R.I. Kaiser, Chem. Rev. 102 (5), 1309–1358 (2002). doi:10.1021/cr970004v.
- [83] J.-D. Chai and M. Head-Gordon, Phys. Chem. Chem. Phys. 10 (44), 6615–6620 (2008). doi:10.1039/b810189b.
- [84] G. Knizia, T.B. Adler and H.-J. Werner, J. Chem. Phys. 130 (5), 054104 (2009). doi:10.1063/1.3054300.
- [85] T.B. Adler, G. Knizia and H.-J. Werner, J. Chem. Phys. 127 (22), 221106 (2007). doi:10.1063/1.2817618.
- [86] T.H. Dunning Jr., J. Chem. Phys. 90 (2), 1007–1023 (1989). doi:10.1063/1.456153.
- [87] J. Zhang and E.F. Valeev, J. Chem. Theory Comput. 8 (9), 3175–3186 (2012). doi:10.1021/ct3005547.
- [88] M.J. Frisch, G.W. Trucks, H.B. Schlegel, G.E. Scuseria, M.A. Robb, J.R. Cheeseman, G. Scalmani, V. Barone, B. Mennucci, G.A. Petersson, H. Nakatsuji, M. Caricato, X. Li, H.P. Hratchian, A.F. Izmaylov, J. Bloino, G. Zheng, J.L. Sonnenberg, M. Hada, M. Ehara, K. Toyota, R. Fukuda, J. Hasegawa, M. Ishida, T. Nakajima, Y. Honda, O. Kitao, H. Nakai, T. Vreven, J.A. Montgomery, J.E. Peralta, F. Ogliaro, M. Bearpark, J.J. Heyd, E. Brothers, K.N. Kudin, V.N. Staroverov, R. Kobayashi, J. Normand, K. Raghavachari, A. Rendell, J.C. Burant, S.S. Ivengar, J. Tomasi, M. Cossi, N. Rega, I.M. Millam, M. Klene, J.E. Knox, J.B. Cross, V. Bakken, C. Adamo, J. Jaramillo, R. Gomperts, R.E. Stratmann, O. Yazyev, A.J. Austin, R. Cammi, C. Pomelli, J.W. Ochterski, R.L. Martin, K. Morokuma, V.G. Zakrzewski, G.A. Voth, P. Salvador, J.J. Dannenberg, S. Dapprich, A.D. Daniels, Ö. Farkas, J.B. Foresman, J.V. Ortiz, J. Cioslowski, D.J. Fox, Gaussian 09, Version A.1., (Gaussian Inc., Wallingford CT, 2009).
- [89] H.-J. Werner, P.J. Knowles, R. Lindh, F.R. Manby, M. Schütz, P. Celani, T. Korona, G. Rauhut, R.D. Amos, A. Bernhardsson, *MOLPRO*, Version 2010.1, A Package of Ab Initio Programs, (University of Cardiff, Cardiff, UK, 2010).
- [90] P.J. Robinson and K.A. Holbrook, Unimolecular Reactions (John Wiley and Sons, New York, 1972).
- [91] H. Eyring, S.H. Lin and S.M. Lin, *Basic Chemical Kinetics* (John Wiley and Sons, New York, 1980).
- [92] J.I. Steinfeld, J.S. Francisco and W.L. Hase, *Chemical Kinetics and Dynamics*, 2nd edition (Pearson, Upper Saddle River, NJ, 1998).
- [93] V.V. Kislov, T.L. Nguyen, A.M. Mebel, S.H. Lin and S.C. Smith, J. Chem. Phys. **120** (15), 7008–7017 (2004). doi:10.1063/1.1676275.
- [94] R.I. Kaiser and A.M. Mebel, Chem. Soc. Rev. 41 (16), 5490 (2012). doi:10.1039/c2cs35068h.
- [95] R.D. Levine, *Molecular Reaction Dynamics* (Cambridge University Press, Cambridge, UK, 2005).
- [96] W.B. Miller, S.A. Safron and D.R. Herschbach, Discuss. Faraday Soc. 44, 108 (1967). doi:10.1039/df9674400108.

- [97] N. Balucani, O. Asvany, R.-I. Kaiser and Y. Osamura, J. Phys. Chem. A. 106 (17), 4301–4311 (2002). doi:10.1021 /jp0116104.
- [98] N. Marcelino, J. Cernicharo, M. Agúndez, E. Roueff, M. Gerin, J. Martín-Pintado, R. Mauersberger and C. Thum, Astrophys. J. 665 (2), L127–L130 (2007). doi:10.1086/521398.
- [99] G.A. Blake, E.C. Sutton, C.R. Masson and T.G. Phillips, Astrophys. J. 315, 621 (1987). doi:10.1086/165 165.
- [100] J. Cernicharo, A.M. Heras, A.G.G.M. Tielens, J.R. Pardo,
 F. Herpin, M. Guélin and L.B.F.M. Waters, Astrophys. J. 546 (2), L123–L126 (2001). doi:10.1086/318
 871.